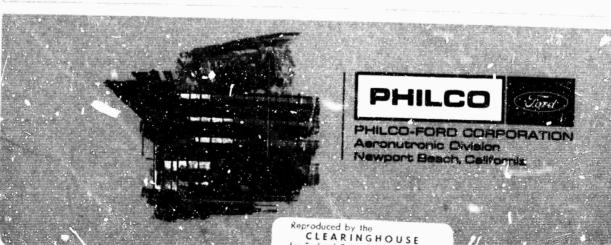
# ABSURPTION BY CO2 BETWEEN 1800 AND 2850 cm<sup>-1</sup> (3.5-5.6 MICRONS)

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ARPA Order No. 237 Amendment #23/I-3-66

15 December 1966

#### SCIENTIFIC REPORT

ABSORPTION BY CO<sub>2</sub> BETWEEN 1800 AND 2850 cm<sup>-1</sup>
(3.5-5.6 Microns)

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#### ABSTRACT

Transmission spectra in the  $1800\text{-}2850~\text{cm}^{-1}$  region have been obtained for more than 100~samples of  $\text{CO}_2$  and  $\text{CO}_2$  mixed with  $\text{N}_2$  and A. The spectral resolution was  $2.5~\text{cm}^{-1}$ . Sample pressures varied from 0.0055~to 742 torr with absorber thicknesses covering the range from 0.081~to 84,400 atm cm. Spectra of several samples at the lower pressures show the effect of Doppler broadening. Measurements in the  $2400\text{-}2560~\text{cm}^{-1}$  region provide information about the absorption by the extreme wings of collision-broadened lines. Replotted transmission spectra and extensive tables of integrated absorptance for 116~samples are included.

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#### SECTION 1

#### INTRODUCTION AND SUMMARY

Absorption and emission by CO<sub>2</sub> in the 1800-2850 cm<sup>-1</sup> region plays a very important part in the transfer of heat in the atmospheres of the earth and other planets. Because of many very strong lines in this region, there is appreciable absorption by atmospheric paths which are so short or at such low pressures that absorption in other regions of the infrared is almost negligible.

Several quantitative measurements on the absorption in this region have been made previously with low resolution for the purpose of determining the relationship between the integrated absorptance  $\int A(\nu) d\nu$  and the parameters, absorber thickness and pressure. The present investigation, which was undertaken to supplement the previous work, includes measurements on samples having much greater absorber thicknesses. Therefore, it has been possible to measure  $\mathrm{CO}_2$  absorption in spectral regions where it had not been observed previously. Other samples with long paths and very low pressures have provided data under conditions for which Doppler broadening of the absorption lines is important. Information on the absorption by the extreme wings of the strongest lines has also been obtained from measurements in the 2400-2560 cm<sup>-1</sup> region.

The experimental methods are discussed in Section 2. Section 3 includes spectral curves for 116 samples of  $\mathrm{CO}_2$  alone and  $\mathrm{CO}_2 + \mathrm{N}_2$  as well as a limited discussion of the results. Extensive tables of the integrated absorptance are included in Section 4. Tables of transmittance vers s wavenumber are available from the authors for workers who require them.

Additional measurements with resolution less than 0.5 cm $^{-1}$  will be made in this region by us in the future. The results will be used to identify many of the very weak bands and to determine the contributions of various bands in regions where several of them may overlap.

#### SECTION 2

#### **EXPERIMENTAL**

#### 2.1 INSTRUMENTAL

Samples of CO<sub>2</sub> alone and mixtures of CO<sub>2</sub> with N<sub>2</sub> or A were contained in a multiple-pass absorption cell whose base length is approximately 29 meters. The cell was used at 4, 8, 16, and 32 passes, giving path lengths of 121, 237, 469, and 933 meters, respectively. Radiation from a Nernst glower traversed the absorption cell and formed an image of the source on the slit of a Perkin Elmer Model 112 spectrometer which employed an LiF prism and a thermocouple detector. While a spectrum was being scanned, the spectrometer slits were adjusted continuously by a string cam which coupled the slit micrometer to the Littrow screw that rotated the prism. The cam, which was designed and built in our laboratory, adjusted the slits so that the signal from the detector was approximately constant while scanning a spectrum with the absorption cell evacuated. The spectral slitwidth was approximately 2.5 cm<sup>-1</sup>.

The monochromator was flushed with dry  $N_2$ , and the remainder of the optical path outside the absorption cell was contained in vacuum tanks in order to eliminate absorption by atmospheric gases. Wavenumber calibration was obtained from  $H_2O$ ,  $CO_2$ ,  $CH_4$ ,  $N_2O$  and CO absorption lines whose positions are known. Details of the multiple-pass cell and the spectrometer have been described previously.  $^3$ ,

#### 2.2 SAMPLING PROCEDURE

The gases used for samples were obtained from commercial cylinders. The  $N_2$  was high-purity dry grade with less than 10 parts  $H_2O$  per million,

and the CO<sub>2</sub> contained traces of  $\rm H_2O$  and CO. It is probably safe to assume that all the isotopes were present in their natural abundances (C<sup>12</sup>, 98.9%; C<sup>13</sup>, 1.1%; C<sup>16</sup>, 99.76%; O<sup>17</sup>, 0.04%; O<sup>18</sup>, 0.20%; h, 99.9844%; D, 9.0156%).

An Hg manometer was used to measure pressures in the range from 50 torr to 1 atm; a manometer containing a special oil was used for pressures between 1.5 and 50 torr. Some of the lower pressures were measured by a McLeod gauge; other pressures of pure CO2 which were too low to measure accurately with any of the gauges were determined by expanding CO, into the cell from a cylinder filled to a pressure that could be measured accurately. The volume of the cylinder was approximately 0.001 times that of the absorption cell. The ratio of the pressure in the cylinder to the resulting pressure in the cell was determined by using enough CO2 that the pressure in the cell was several torr, which was high enough to be measured accurately. We then assumed that the ratio of pressures was the same at lower pressures. When the initial pressure in the cylinder was greater than approximately 1 atm, it was necessary to account for the non-linearity in the relation between CO2 density and pressure. From the Van der Waal's constants for CO2, we can show that the density is proportional to p(1 + 0.005p) if p, the pressure in atm. is less than approximately 15.

Adsorption of CO2 on the walls of the cylinder and the absorption cell probably gives rise to the greatest uncertainty in determining CO2 pressures by the expansion method. If the percent adsorbed was independent of pressure, very little error was introduced. However, it seems possible that a greater percentage of gas would be adsorbed when it is first added until a film is formed on the surface; after this, the percentage adsorbed would decrease as the pressure increases. No measurements were made to determine if such a saturation phenomenon occurred in our system. But if it did occur at pressures less than approximately 0.1 torr, the values we used for very low CO2 pressures are probably too high. We compared the integrated absorptance of a few samples of  $CO_2 + N_2$  in which the  $CO_2$ pressure was determined by expansion to some previous results for samples with shorter paths and higher  ${
m CO_2}$  pressures which could be measured accurately. The integrated absorptance of the earlier samples was usually slightly greater than that of present samples having the same absorber thickness and equivalent pressure. Therefore, it seems likely that there were small systematic errors, possibly due to adsorption, in the pressures we determined by the expansion method.

In view of the above discussion, the quoted values of  ${\rm CO_2}$  pressures below 0.1 torr are probably less than 8 percent too high or less than 2 percent too low.

Mixtures of  ${\rm CO_2}$  +  ${\rm N_2}$  or  ${\rm CO_2}$  + A were formed by adding the  ${\rm N_2}$  or A to the cell after the  ${\rm CO_2}$  was introduced. Fans installed in the cell were used

to mix the gases. Several different samples, each at a different total pressure, were formed from the same C)<sub>2</sub>. The same mixture was also frequently investigated at four different path lengths: 121, 237, 469, and 933 meters.

The absorber thickness u was calculated by the use of the fo'lowing equation.

$$u(atm cm)_{STP} = (1 + 0.005p) p L 273/296,$$
 (2-1)

where L is the geometrical path length in cm and p is the partial pressure of  $\mathrm{CO}_2$  in atm. The term (273/296) accounts for the difference in density between standard temperature  $(273^\circ\mathrm{K})$  and room temperature  $(296^\circ\mathrm{K})$  at which the measurements were made. The quantity (1+0.005p), which accounts for the non-linearity in the relation between the density of  $\mathrm{CO}_2$  and its pressure, is negligible except for pressures greater than approximately 1 atm. It could be neglected for samples included in the present study, but it has been included in a computer program used to calculate sample parameters for pressures as high as 15 atm.

When working with mixtures of  ${\rm CO_2}$  +  ${\rm N_2}$ , it is convenient to use an equivalent pressure  ${\rm P_e}$  which is proportional to the half-width of the absorption lines, regardless of the composition of the mixture. We have found that such an equivalent pressure is given by

$$P_p = 1.3 p + (P - p),$$
 (2-2)

where P is the total pressure, and p is the partial pressure of  ${\rm CO}_2$ . It is noted that P<sub>e</sub> approaches P for a very dilute mixture of  ${\rm CO}_2$  in N<sub>2</sub> (p << P).

Table 2-1 includes the parameters for 116 samples of  $\mathrm{CO}_2$  and  $\mathrm{CO}_2$  +  $\mathrm{N}_2$ . The  $\mathrm{CO}_2$  partial pressure p, the total pressure P, and the equivalent pressure  $\mathrm{P}_{\mathrm{e}}$  are given in torr and in atm. Also included are references to the transmittance curves and the integrated absorptance tables. Samples of  $\mathrm{CO}_2$  + A which are discussed in Section 3.3 were scanned only over the region above 2400 cm<sup>-1</sup> and are not included in Table 2-1.

#### 2.3 RECORDING AND REDUCTION OF DATA

A spectrum of each sample was scanned over a sufficiently wide region that there was essentially no absorption at the starting and end points. Spectral curves called background curves were scanned over the same spectral regions with the cell evacuated. The shapes of the background curves varied with the number of passes of the cell because of the variation in reflectivity with wavenumber. Therefore, it was necessary to scan background curves at the same paths as those used for the samples.

TABLE 2-1
SAMPLE PARAMETERS

Sam. No.	р	P	Pe	р	P	Pe
	torr	torr	torr	atm	atm	atm
1	742	742	969	0.976	0.976	C,275
2	742	742	969	0.976	0.976	0.275
3	208	208	2 <b>7</b> 1	0.274	0.274	0,35€
4	742	742	969	0.976	0.976	1.28
5	101	101	131	0.133	0.133	0.173
6	208	208	271	0.274	0.274	0.356
7	208	/40	803	0.274	0.974	0.0563
8	51.5	51.5	67.0	0.0678	0.0678	0.0881
9	101	101	131	0.133	0.133	0.173
10	208	208	271	0.274	0.274	0.356
11	208	740	503	0.274	0.974	1.06
12	51.5	51.5	67.0	0.0678	0.0678	0.0881
13	208	208	271	0.274	0.274	0.356
14	208	740	803	0.274	0.974	1.06
15	26.8	26.8	34.8	0.0353	0.0353	0.0459
16	26.8	229	237	C.0353	0.301	0.312
17	26.7	26.7	34.7	0.0351	0.0351	0.0457
18	۷.7	229	237	0.0351	0.301	0.312
19	3.20	3.20	4.16	0.00421	0.00421	0.00547
20	26.7	26.7	34.7	0.0351	0.0351	0.0457
21	26.7	229	237	0.0351	0.301	0.312
22	3.20	3.20	4.16	0.00421	0.00421	0.90547
23	3.20	10.9	11.9	0.00421	0.0143	0.0156
24	3.20	32.9	33.9	0.00421	0.0433	0.0446
25	3.20	103	104	0.00421	0.136	0.137
26	0.80	0.80	1.04	0.00105	0.00105	0.00137
27	3.20	3.20	4.16	<b>0.00421</b>	0.00421	0.00547
28	3.20	10.9	11.9	0.00421	G.0143	0.0156
29	3.20	32.5	33.5	0.00421	0.0428	0.0440
30	3.20	103	104	0.00421	0.136	0.137
31	0.80	0.80	1.04	0.00105	0.00105	0.00137
32	3.20	3.20	4.16	0.00421	0.00421	0.00547
33	3.20	10.9	11.9	0.00421	0.0143	0.0156
34	3.20	32.5	33.5	0.00421	0.0428	0.0440
35	3.20	103	104	0.00421	0.136	0.137
36	0.400	0.400	0.520	0.000526	0.000526	0.000684
37	0.400	1.00	1.12	0.000526	0.00132	0.00147
38	0.400	3.20	3.32	0.000526	0.00421	0.00437
39	0.400	15.0	15.1	0.000526	0.0197	0.0199
40	0.400	100	_100.1	0.000526	0.132	0.132

TABLE 2-1 (cont.)

Sam. No.	L Path	u atm cm STP	Fig. in which spectral curve	Tables of integrated
	ın		appears	absorptance
1	933	84,400	3-1	4-1
2	469	42,400	3-1	4-1
3	933	23,600	3-1	4-1
4	237	21,400	3-1	4-1
5	933	11,400	3-1	4-1
6	+69	11,900	3-1	4-1
7	469	11,900	3-1	4-1
8	933	5, 830	3-1	4-1
9	469	5,750	3-1	4-1
10	237	5,990	3-1	4-1
11	237	5,990	3-2	4-1
12	469	2,930	3-2	4-1
13	121	3,060	3-2	4-1
14	121	3,060	3-2	4-1
15	469	1,530	3-2	4-1
16	469	1,530	3-2	4-1
17	237	768	3-2	4-2
18	237	768	3-2	4-2
19	933	362	3-2	4-2
20	121	392	3-2	4-2
21	121	392	3-3	4-2
22	469	182	3-3	4-2
23	469	182	3-3	4-2
24	469	182	3-3	4-2
25	469	182	3-3	4-2
26	933	90.6	3-3	4-2
27	237	92.0	3-3	4-2
28	237	92.0	3-3	4-2
29	237	92.0	3-3	4-2
30	237	92.0	3-3	4-2
31	469	45.5	3-3	4-2
32	121	47.0	3-3	4-2
33	121	47.0	3-3	4-2
34	121	47.0	3-3	4-2
35	121	47.0	3-3	4-2
36	469	22.8	3-4	4-3
37	469	22.8	3-4	4-3
38	469	22.8	3-4	4-3
39	469	22.8	3-4	4-3
40	469	22.8	3-4	4-3

TABLE 2-1 SAMPLE PARAMETERS

Sam. No.	p	P	Pe	P	P	P <sub>c</sub>
140.						
	torr	to~r	torr	atm	atm	atm
41	0.100	0.100	0.130	0.000132	0.000132	0.000171
42	0.200	0.200	0.260	0.000263	0.00 263	0.000342
43	0.400	0.400	0.520	0.000526	0.000526	0.000684
44	0.400	1.00	1.12	0.000526	0.00132	0.00147
45	0.400	3.20	3.32	0.000526	0.00421	0.00437
46	0.400	15.0	15.1	0.000526	0.0197	0.0199
47	0.400	100.0	100.1	0.000526	0.132	0.132
48	0.051	0.051	0.066	0.000067	0.000067	0.000087
49	0.100	0.100	0.130	0.000132	0.600132	0.000171
50	0.200	0.200	0.260	0.000263	0.000263	0.000342
53	0.400	0.400	0.520	0.000526	0.000526	0.000684
52	0.400	1.00	1.12	0.000526	0.00132	0.00147
53	0.400	3.20	3.32	0.000526	0.00421	0.00437
54	0.400	15.0	15.1	0.000526	0.0197	0.0199
55	0.400	100.0	100.1	0.000526	0.132	0.132
5 <b>6</b>	0.025	0.025	0.033	0.000033	0.000033	0.000043
57	0.025	0.054	0.062	0.000033	0.000071	0.000081
58	0.025	0.114	0.122	0.000033	0.000150	0.000160
59	0.025	0.294	0.302	0.000033	0.000387	0.000397
60	0.025	0.723	0.731	0.000033	0.000951	0.000961
61	0.025	1.88	1.89	0.000033	0.00247	0.00248
62	0.025	5.25	5.26	0.000033	0.00691	0.00692
63	0.025	14.5	14.5	0.000033	0.0191	0.0191
64	0.025	39.0	39.0	0.000C33	0.0513	0.0513
65	0.025	100.0	100.0	0.000033	0.132	0.132
66	0.012	0.012	0.016	0.000016	0.000016	0.006021
67	0.025	0.025	0.033	0.000033	0.000033	0.000043
68	0.025	0.054	0.062	0.000033	0.000071	0.000081
69	C.025	0.114	0.122	0.000333	0.000150	0.000160
70	0.025	0.294	0.302	0.000033	0.500387	0.000397
71	0.025	0.723	0.751	0.000033	0.000951	0.000961
72	0.025	1.88	1.89	0.000033	0.00247	0.0248
73	0.025	5.25	5.26	0.000033	0.00691	0.00692
74	0.025	14.5	14.5	0.000033	0.0191	0.0191
75	0.025	39.0	39.0	0.000033	0.0513	0.0513
76	0.025	100.0	100.0	0.000033	0.132	0.132
77	0.0055	0.0055	0.0072	0.00000	0.0000372	0.0000094
78	0.012	0.012	0.016	0.000016	0.000016	0.000021
79	0.025	0.025	0.033	0.000033	0.000033	0.000043
80	0.025	0.054	0.062	0.000033	0.000071	0.000081

TABLE 2-1 (cont.)

Sam. No.	L Path m	u atm cm STP	Fig. in which spectral curve appears	Tables of integrated absorptance
41	933	11.3	3-4	4-3
42	469	11.4	3-4	4-3
43	237	11.5	3-4	4-3
44	237	11.5	3-4	4-3
45	237	11.5	3-4	4-3
46	237	11.5	3-4	4-3
47	237	11.5	3-4	4-3
48	933	5.8	3-4	4-3
49	469	5.69	3-4	4-3 4-3
50	237	5.75	3-4 3-4	4-3
51	121	5.87	3-4	4-3
52	121	5.87	3-4	4-3
53	121	5.87	3-4	4-3 4-3
54	121	5.87		
55	121	5.87	3-4 3-4	4-3 4-3
56	933	2.8	3-4	4-4
57	933	2.8	3-4	4-4
58	933	2,8	3-4	4-4
59	933	2.8	3-4	4-4
60	933	2.8	3-4	4-4
61	933	2.8	3-4	4-4
62	933	2.8	3-4	4-4
63	933	2.8	3-4	4-4
64	933	2.8	3-4	4-4
65	933	2.8	3-4	4-4
66	933	1.4	3 <b>-</b> 5	4-4
67	469	1.4	3-5	4-4
68	469	1.4		
69	469	1.4	3-5 3-5	4-4 4-4
70	469	1.4	3 <b>-</b> 5	4-4 4-4
71	469	1.4	3-5	4-4
72	469	1.4	3-5	4-4
73	469	1.4	3-5	4-4
74	469	1.4	3-5	4-4
75	469	1.4	3-5	4-4
76	469	1.4	3-5	4-4
77	933	0.62	3-5	4-5
78	469	0.68	3-5	4-5
79	237	0.72	3-5	4-5
80	237	0.72	3-5	4-5

TAE: E 2-1
CAMPLE PARAMETERS

Sam.	p	P	Pe	р	P	P_
No.	-		е	-		e
	torr	torr	torr	atm	atm	atm
81	0.025	0.114	0.122	0.000033	0.000150	0.000160
82	0.025	0.294	0.302	0.000033	0.000387	0.000397
83	0.025	0.723	0.731	0.000033	0.000951	0.000961
84	0.025	1.88	1.89	0.000033	0.00247	0.00248
85	0.025	5.25	5.26	0.000033	0.00691	0.00692
	0.023	3.23	3.20	0.000033	0.00071	0,000,2
86	0.025	14.5	14.5	0.000033	0.0191	0.0191
87	0.025	39.0	39.0	0.000033	0.0513	0.0513
88	0.025	100.0	100.0	0.000033	0.132	0.132
89	0.0055	0.0055	0.0072	0.0000072	0.0000072	U.0000094
90	0.012	0.012	0.016	0.000016	0.000016	0.000021
_		<b></b>				- •
91	0.025	0.025	0.033	0.000033	0.000033	0.000043
92	0.025	0.054	0.062	0.000033	0.000071	0.000081
93	0.025	0.114	0.122	0.000033	0.000150	0.000160
94	0.025	0.294	0.302	0.000033	0.000387	0.000397
95	0.025	0.723	0.731	0.000033	0.000951	0.000961
96	0.025	1.88	1.89	0.000033	0.00247	0.00248
97	0.025	5.25	5.26	0.000033	0.00691	0.00692
98	0.025	14.5	14.5	0.000033	0.0191	0.0191
99	0.025	39.0	39.0	0.000033	0.0513	0.0513
1⊕0	0.025	100.0	100.0	0.000033	0.132	0.132
17.0	0.023	100.0	100.0	0.00,005	0.132	3.132
101	0.0055	0.0055	0.0072	0.0000072	0.0000072	0.0000094
102	0.012	0.012	0.016	0.000016	0.000016	0.000021
103	0.012	0.126	0.130	0.000016	0.000166	0.000171
104	0.012	0.384	0.388	0.000016	0.000505	0.000510
105	0.012	1.20	2.20	0.000016	0.00158	0.00158
106	0.012	3.45	3.45	0.000016	0.00454	0.00454
107	0.012	10.2	10.2	0.000016	0.0134	0.0134
108	0.012	32.5	32.5	0.000016	0.0428	0.0428
109	0.012	100.4	100.4	0.000016	0.132	0.132
110	0.0055	د0.005	0.0072	0.0000072	0.0000072	0.0000094
111	0 0055	0 107	0 100	0.0012070	0.000050	0.0000(1
111 112	0.0055	0.197 0.600	0.199	0.0CJ0072	0.000259	0.000261
	0.0055	1.89	0.602	0.0000072	0.000789	0.000792
113	0.0055	8.51	1.89	0.0000072	0.00249	0.00249
114	0.0055 0.0055	79.5	8.51 29.5	0.0000072	0.0112	0.0112
115	0.0000	. 7. 3	47.3	0.0000072	0.0388	0.0388
1.6	0.0055	102	102	0.0000072	0.134	0.134

TABLE 2-1 (cont.)

Sam. No.			Fig. in which spectral curve appears	Tables of integrated absorptance
	m			
81	237	0.72	3-5	4 <b>-</b> 5
82	237	0.72	3 <b>-</b> 5	4~5
83	237	0.72	3 <b>-</b> 5	4 <b>-</b> 5
84	237	0.72	3 <b>-</b> 5	4 <b>-</b> 5
85	237	0.72	3-5	4 <b>-</b> 5
86	237	Ü.72	3 <b>-</b> 5	4 <b>-</b> 5
87	237	0.72	3-5	4 <b>-</b> 5
38	237	0.72	3-5	4 <b>-</b> 5
89	469	9.31	3 <b>-</b> 5	<b>4</b> <i>∽</i> 5
90	237	0.35	3-5	4 <b>-</b> 5
91	121	0.37	3-5	4 <b>-</b> 5
92	121	0.37	3-5	4 <b>-</b> 5
93	121	0.37	3-5	4-5
94	121	0.37	3-5	4-5
95	121	0.37	3-5	4-5
96	121	0.37	3-6	4 <b>-</b> 5
97	121	0.37	3-6	4-5
98	121	0.37	3-6	4 <b>-</b> 5
99	121	0.37	3-6	4 <b>-</b> 5
100	121	0.37	3-6	4-5
101	237	0.16	3-6	4-6
102	121	0.18	3-6	4-6
103	121	0.18	3 <b>-</b> 6	4-6
104	121	0.18	3-6	4-6
105	121	0.18	3-6	4-6
106	121	0.18	3-6	4-6
107	121	0.13	3 <b>-</b> 6	4 <b>-</b> 6
108	121	0.18	3 <b>-</b> 6	4-6
100	121	0.18	3-6	4-6
110	121	0.081	3-6	4 <b>-</b> 6
111	121	0.081	3-6	4-6
112	121	0.081	3-6	4 <b>-</b> 6
113	121	0.081	3-6	4 <b>-</b> 6
113	121	0.081	3-6	4 <b>-</b> 6
115	121	0.081	3-6	4 <b>-</b> 6
116	121	0.081_	3-6	4-6

Each spectrum was examined and compared with others as a check for consistency. Small corrections were made to account for spurious deflections and for absorption by  $H_2O$  and CO impurities in the sample. The transmittance was determined from the ratio of the deflection on the sample curve to the deflection on the background curve at the same wavenumber. Each spectral curve then was replotted and digitized by the method describe! previously. Pairs of values related to transmittance and wavenumber were punched on IBM cards which served as input for a computer program used to calculate transmittance and integrated absorptance as a function of wavenumber. The replotted spectra are shown in Section 3 and tables of integrated absorptance appear in Section 4.

#### SECTION 3

#### RESULTS AND DISCUSSION

#### 3.1 TRANSMISSION SPECTRA

Curves of transmittance versus wavenumber are shown in Figs. 3-1 through 3-6 for the 116 samples of  $\rm CO_2$  and  $\rm CO_2$  + N<sub>2</sub> listed in Table 2-1. The curves were replotted from the original curves obtained with a spectral resolution of approximately 2.5 cm<sup>-1</sup>. Small corrections were made to account for absorption by CO near 2140 cm<sup>-1</sup> and by H<sub>2</sub>O from 2800 to 2870 cm<sup>-1</sup> and from 1815 to 1870 cm<sup>-1</sup>.

Table 3-1 includes a list of absorption bands expected in this region. Evidence of many of them can be seen in the transmission spectra, although most of the absorption is due to the very strong  $00^01$  band and two medium strength bands,  $11^{10}$  and  $03^{10}$ . Features of several of the bands listed in Table 3-1, as well as others not listed, can probably be identified in spectra with higher resolution which we plan to obtain.

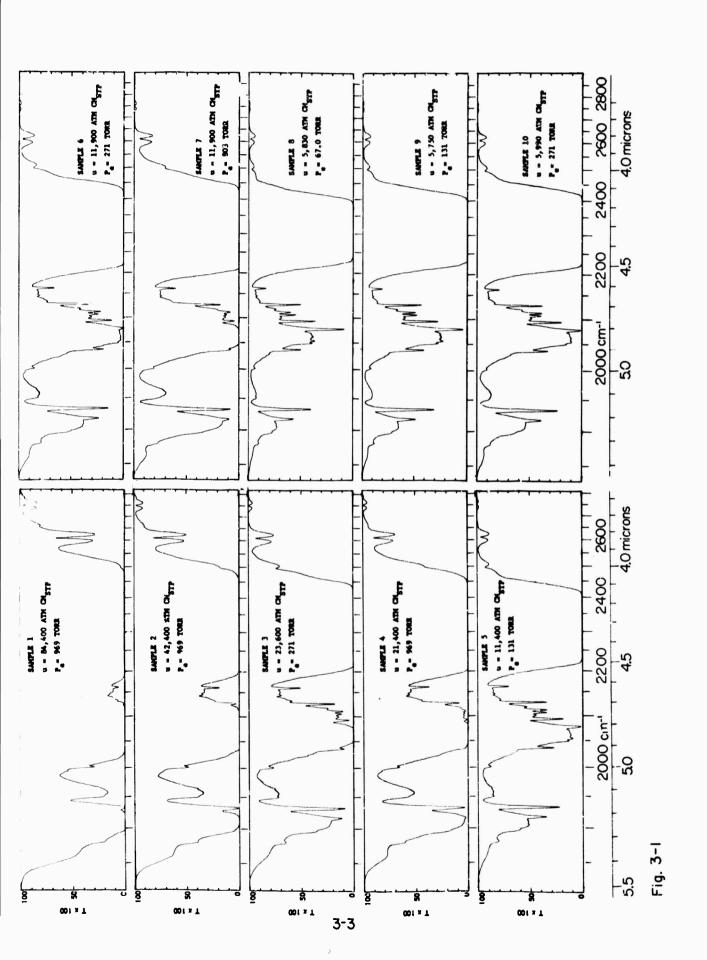
TABLE 3-1 CO<sub>2</sub> ABSORPTION BANDS BETWEEN 1800 AND 2800 cm<sup>-1</sup>

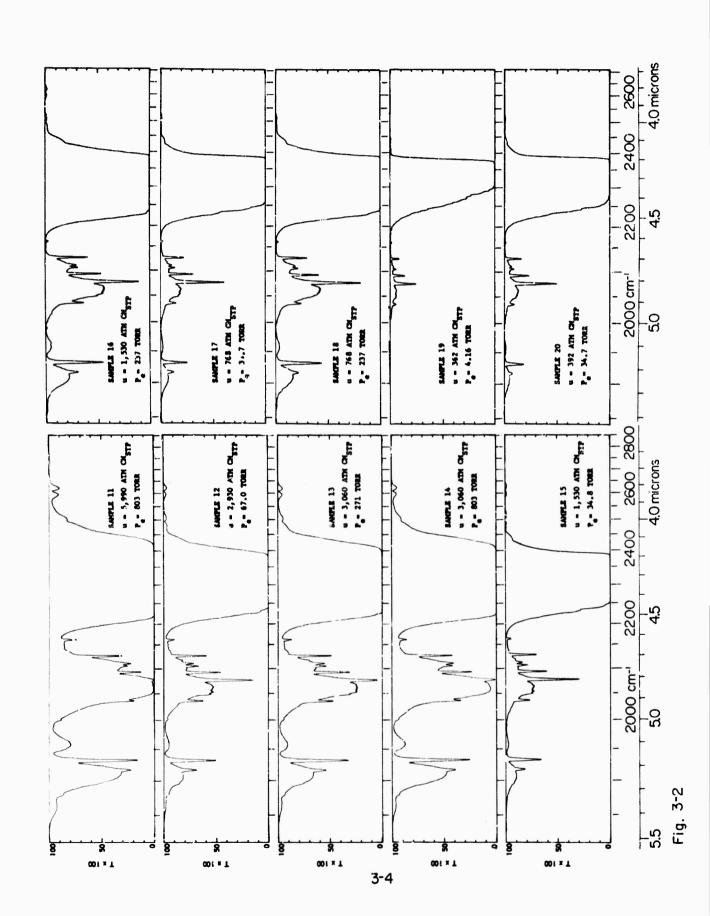
Band Center	Upper Level	Lower Level*	Molecular Species <sup>+</sup>
1846.29	05 <sup>1</sup> 0	02 <sup>2</sup> 0	
1886. н	0400	0110	
1896.00	0510	0200	
1917.67	04 <sup>2</sup> 0	0110	
1932.5 н	03 <sup>1</sup> 0	01 0	
1932.5 H		•	
2003.5	1200	01 <sup>1</sup> 0	
2004.01	13 <sup>1</sup> 0	02 <sup>2</sup> 0	
2053.72	13 <sup>1</sup> 0	02 <sup>0</sup> 0	
2076.5 н	11 <sup>1</sup> 0		
2094.	12 <sup>2</sup> 0	01 <sup>1</sup> 0	
2137. н	2000	01 <sup>1</sup> 0	
2165.30	21 <sup>1</sup> 0	02 <sup>2</sup> 0	
2215.01	21 <sup>1</sup> 0	02 <sup>2</sup> 0	
2327.48	02 <sup>0</sup> 1	02 <sup>0</sup> 0	
2336.66	01 <sup>1</sup> 1	01 <sup>1</sup> 0	
2349.3 н	0001		
2429.41	10 <sup>0</sup> 1		
2500.42	04 <sup>0</sup> 0		$c^{12}0^{16}0^{18}$
2548.33	04 <sup>0</sup> 0 PI		
2614.24	1200		$c^{12}0^{16}0^{18}$
2670.90	12 <sup>0</sup> 0 PI		
2757.04	20 <sup>0</sup> 0		$c^{12}0^{16}0^{18}$
2797.02	20 <sup>0</sup> 0 PI		

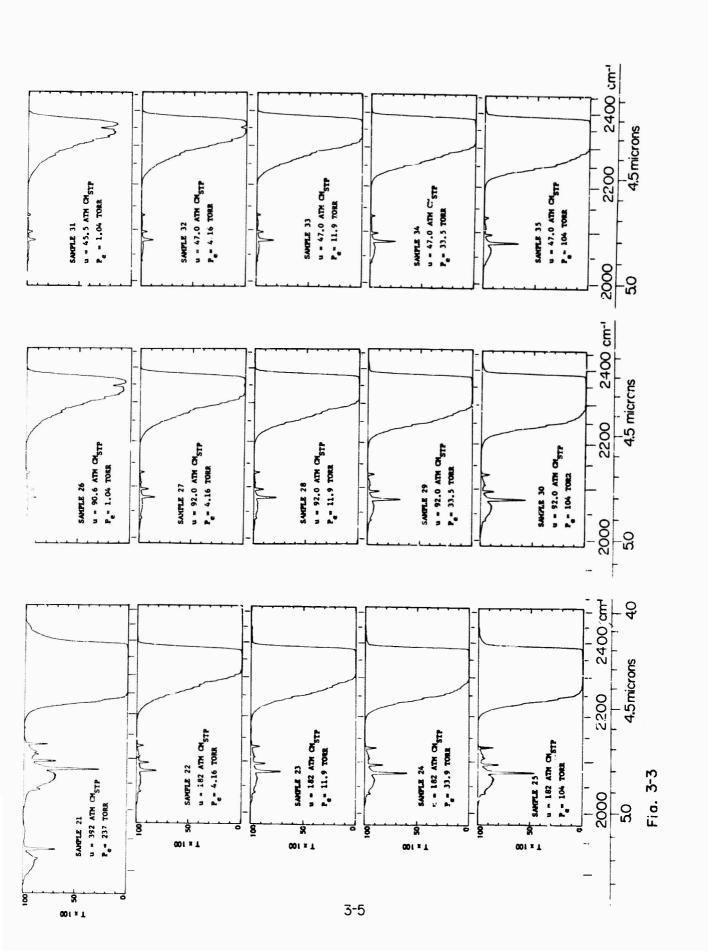
H denotes that the position of the band center is from Herzberg<sup>5</sup>; all others were calculated from energy levels given by Stull, Wyatt and Plass<sup>6</sup>.

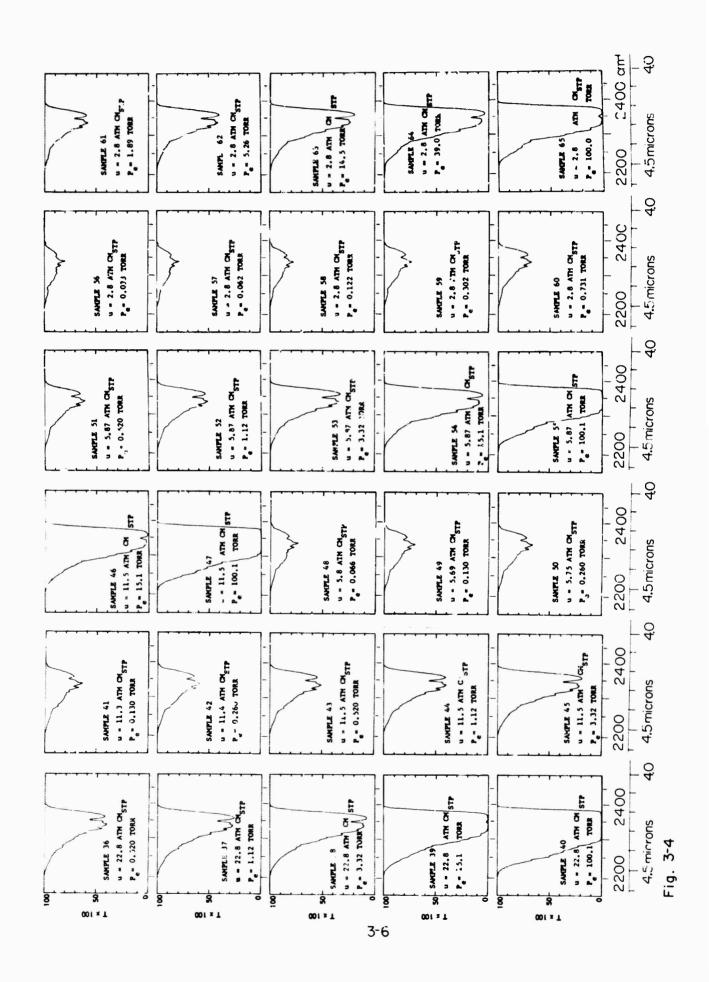
Lower level is  $00^0$ 0 unless indicated otherwise. All species are the  $C^{12}0^{16}0^{16}$  molecule except as noted.

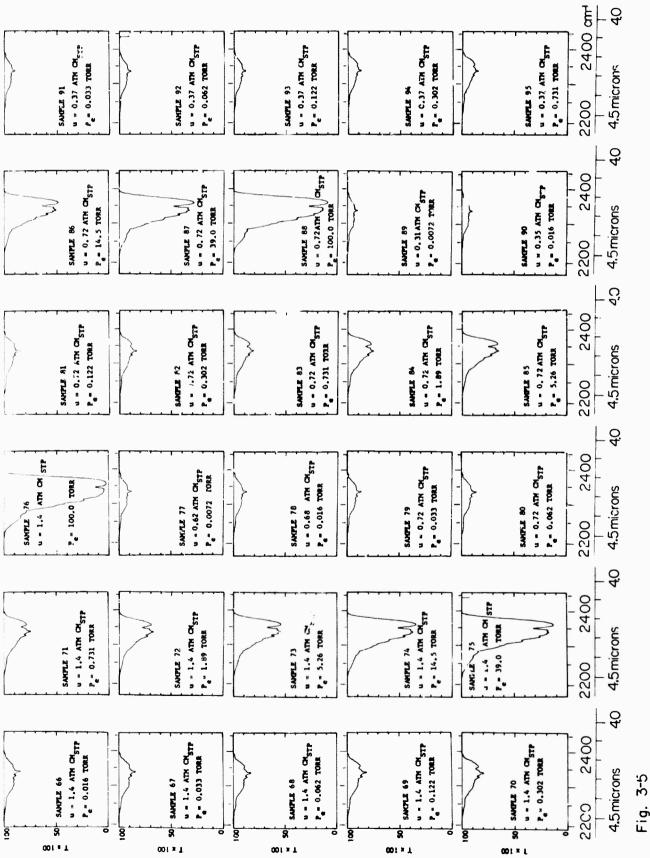
PI denotes pressure-induced bands.

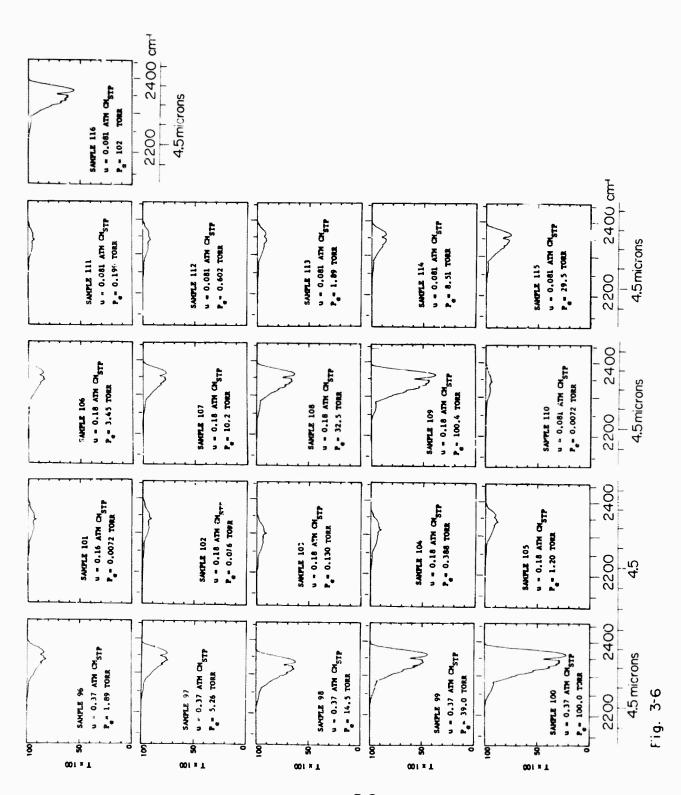












#### 3.2 INTEGRATED ABSORPTANCE

The integrated absorptance over the region from 2190 to 2425 cm<sup>-1</sup> is plotted versus  $P_e$  on log-log scales in Fig. 3-7 for Samples 36 through 116. Each solid curve corresponds to the value of absorber thickness indicated. The broken curve with slope of 0.5 has been included for comparison with the other curves. The integrated absorptance of a band composed of non-overlapping strong lines having the Lorentz line shape is proportional to  $(uP_e)^{0.5}$ . (A strong line is essentially opaque over a region a few times as wide as the width of the line.) The  $(uP_e)^{0.5}$  dependence would give rise to a curve of slope 0.5 on the log-log plot in Fig. 3-7. We see that the slopes of several of the curves are slightly less than 0.5 for pressures between 10 and 100 torr. The deviation from the  $(uP_e)^{0.5}$  relationship for pressures greater than 10 torr is due to overlapping of the lines and the presence of weak lines. The effect of overlapping is particularly important for the larger values of u.

The slopes of the curves representing the smaller values of absorber thickness are seen to decrease with decreasing pressure. The increased absorptance at low pressure is due to the Doppler broadening of the absorption lines. The Lorentz line shape, which is a good approximation to collision-broadened lines, is quite different from the pressure independent Doppler line shape. The absorption coefficient in the wings of a Doppler shaped line decreases much more rapidly with the distance from the center than does a Lorentz line. Therefore, under certain conditions, the absorption in the wings of a line is due to collisi on broadening, while Doppler broadening dominates near the line center. Essentially all the absorption by a low pressure sample with very small absorber thickness occurs near the line center; therefore, its integrated absorptance is independent of pressure. However, in the case of a low pressure sample with intermediate absorber thickness, there is appreciable absorption in the wings of the lines where collision broadening is dominant. Therefore, the integrated absorptance is slightly dependent on pressure. The increasing dependence on  $P_e$  as u increases can be seen by comparing the slopes of the curves in Fig. 3-7 in the region near  $P_e$  = 0.1 to r. Plass<sup>7</sup> has given a theoretical discussion of the absorption by lines in which either Doppler broadening or collision broadening is dominant as well as lines in which both types of broadening make significant contributions.

Figure 3-8 shows the relation between integrated absorptance and absorber thickness for different values of  $P_{\rm e}$ . The curve corresponding to 1000 torr represents data from Burch, Gryvnak, and Williams and is included for comparison. The other curves were cross plotted from the curves in Fig. 3-7. Curves corresponding to absorption by non-overlapping strong lines with the Lorentz shape would be parallel to the comparison line whose slope is 0.5. Segments of the 10 torr and 100 torr curves are

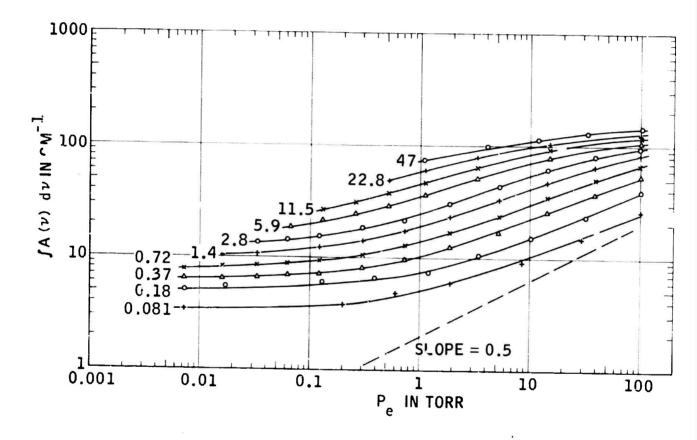


Fig. 3-7 THE INTEGRATED ABSORPTANCE OF THE 2190-2425 cm<sup>-1</sup> REGION VERSUS EQUIVALENT PRESSURE.

Each curve corresponds to the indicated value of absorber thickness in atm  $cm_{\mbox{STP}}$ . The broken line with slope = 0.5 is shown for comparison.

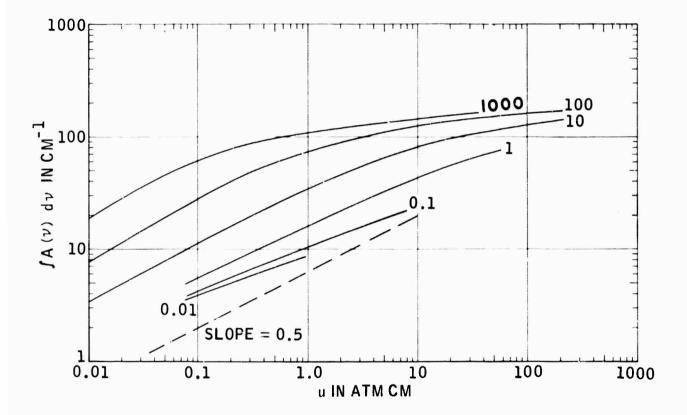


Fig. 3-8 THE INTEGRATED ABSORPTANCE OF THE 2190-2425 cm<sup>-1</sup> REGION VERSUS ABSORBER THICKNESS.

Each curve corresponds to the indicated value of equivalent pressure in torr. The broken line with slope -0.5 is shown for comparison.

seen to be nearly parallel to the comparison line, indicating that integrated absorptance is approximately proportional to  $u^{0.5}$  for the values of u and P<sub>p</sub> represented.

The relation between the integrated absorptance and the parameter uP<sub>e</sub> is shown in Fig. 3-9. The integrated absorptance can be expressed as a function of this convenient parameter when the absorption is primarily due to strong lines with the Lorentz shape. Under this condition, all the curves corresponding to different pressures coincide. Although none of the curves coincide, except when the absorption is nearly complete throughout much of the band, the 1, 10, and 100 torr curves occur near each other for uP<sub>e</sub> greater than approximately 10 atm cm<sub>STP</sub> torr. For smaller values of uP<sub>e</sub> at lower pressures, the curves are separated because of Doppler broadening, indicating that the integrated absorptance cannot be related to the single variable uP<sub>e</sub>.

#### 3.3 ABSORPTION BETWEEN 2400 AND 2580 cm<sup>-1</sup>

A few very weak isotopic bands and two pressure-induced bands occur between  $2^{6}00$  and  $2580~\rm{cm}^{-1}$ , but most of the absorption in this region is due to the extreme wings of the very . ong lines of the  $00^{0}$ l band. The centers of all the lines of this band are confined to the region below the band head near  $2400~\rm{cm}^{-1}$ . We were able to account for the isotopic and pressure-induced bands in the  $2400-2580~\rm{cm}^{-1}$  region and to determine the amount of absorption by the wings of the strong lines. From the results we were able to derive curves from which the absorptance due to the wings of the strong lines can be determined for samples of  $CO_2$ ,  $CO_2 + N_2$ , or  $CO_2 + A$ .

The transmittance  $T(\nu)$  at wavenumber  $\nu$  is related to absorber thickness u and absorption coefficient  $K(\nu)$  according to the following equation.

$$T(v) = \exp \left[-K(v) u\right], \quad \text{or} \quad K(v) = -\frac{1}{u} \ln T(v).$$
 (3-1)

The total absorption coefficient K( $\nu$ ) due to the wings of CO $_2$  lines broadened by CO $_2$  and N $_2$  is given by

$$K(v) = \left[p/p^{\circ}\right] K_{5}^{\circ}(v) + \left[p_{N_{2}}/p^{\circ}\right] K_{N_{2}}^{\circ}(v).$$
 (3-2)

The quantity  $K_s^o(\nu)$  is the self-broadening absorption coefficient which arises from  $CO_2$ - $CO_2$  collisions when the  $CO_2$  pressure is 1 atm. Similarly,  $K_{N_2}^o(\nu)$  is the  $N_2$ -broadening coefficient due to  $CO_2$ - $N_2$  collisions when the

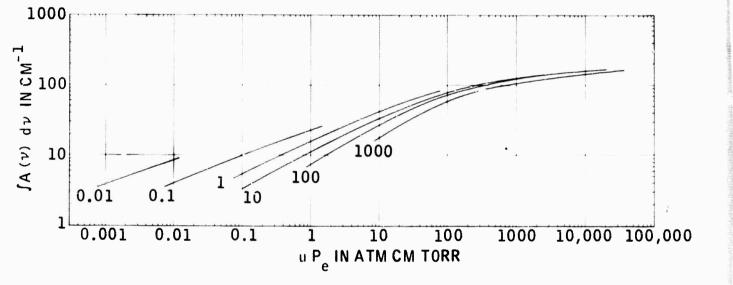


Fig. 3-9 THE INTEGRATED ABSORPTANCE OF THE 2190-2425 cm<sup>-1</sup> REGION VERSUS THE PRODUCT OF ABSORBER THICKNESS AND EQUIVALENT PRESSURE.

Each curve corresponds to the indicated value of equivalent pressure.

 $N_2$  partial pressure is 1 atm. The superscripts (°) denote standard pressure, 1 atm. The partial pressures of  $CO_2$  and  $N_2$  in i m are p and  $P_{N_2}$ , respectively. Equation (3-2) can be used for mixtures of  $CO_2$  plus any non-absorbing broadening gas, such as A, by substituting the appropriate broadening coefficient and partial pressure.

Since no line centers occur in this region, except for those in the very weak bands whose absorption was accounted for, there is no unresolved structure within the 2.5 cm<sup>-1</sup> spectral slitwidth. Therefore, the observed transmittance is a very good approximation to the true transmittance. The absorption coefficient determined from the observed transmittance by the use of Eq. (3-1) also approximates the true coefficient that would be observed with infinite resolution.

Values of the normalized self-broadening coefficient  $K_S^O(\nu)$  were determined from several of the larger samples of pure  $CO_2$  by the use of Eqs. (3-1) and (3-2). These values were then substituted in Eq. (3-2) in order to find values of  $K_N^O(\nu)$  and  $K_A^O(\nu)$  from samples containing these broadening gases. The 2 $\nu$  its are shown in Fig. 3-10, where each of the normalized absorption coefficients is plotted against wavenumber. Points have not been included in the curves at wavenumbers where there is appreciable absorption by the isotopic and pressure-induced bands. Therefore, these curves represent only the contribution of the wings of strong lines whose centers occur below 2400 cm<sup>-1</sup>. Winters, Silverman, and Benedict<sup>8</sup> have made similar measurements in this region. Their work does not extend to wavenumbers as high as ours, but the two sets of results are in good agreement over the region covered by both.

Since the positions, strengths, and widths of the lines are known, it is apparent that considerable information about the shapes of the extreme wings of the lines can be obtained. A report  $^9$  dealing with the shapes of the lines in this region, as well as in the 1.4  $\mu$  and 2.7  $\mu$  regions is being prepared.

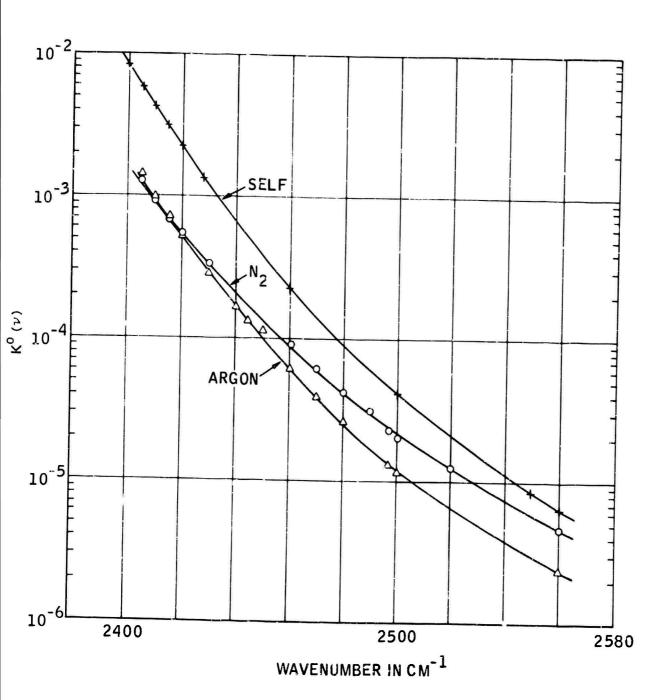


Fig. 3-10 THE NORMALIZED ABSORPTION COEFFICIENT VERSUS WAVENUMBER FOR CO<sub>2</sub> BETWEEN 2400 AND 2580 cm<sup>-1</sup>.

The upper curve corresponds to self-broadened  ${\rm CO}_2$ , i.e., pure  ${\rm CO}_2$ , at 1 atm pressure. The lower two curves correspond to samples of  ${\rm CO}_2$  diluted in the gases indicated at 1 atm. The curves represent only the contribution of the lines whose centers occur below 2400 cm<sup>-1</sup>.

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#### SECTION 4

#### TABLES OF INTEGRATED ABSOLPTANCE

Values of the integrated absorptance  $\int\limits_{V}^{V'} A(v) dv$  for Samples 1 to 116 are shown in Tables 4-1 to 4-6. The sample number and the parameters are listed at the top of each column along with v', the upper limit of integration. The integrated absorptance between any two wavenumbers tabulated is equal to the difference between the values given at those two points.

The integrated absorptance was calculated from values of transmittance which were determined from the spectra at points 1 cm<sup>-1</sup> apart. This interval is sufficiently small that the original spectra can be reconstructed with little loss of structure by plotting the transmittance values and joining the points with straight lines.

The samples included in the various tables are as follows:

San	np1	e No.	<u>Table</u>
1	to	16	4-1
17	to	35	4-2
36	to	55	4-3
56	to	76	4-4
77	to	100	4-5
101	to	116	4-6

					T	able	4-	-I <i>Į</i>	$\Delta(\nu)$	/) dν	•					
San, Bo, p(otm) P <sub>q</sub> (otm) v m)	\$ 9.76 6 10° 5 2.13 6 10° 5 8.44 7 10° 6	2 9.76 6.10 <sup>-1</sup> 9.73 8.10 <sup>-1</sup> 6.26 8.10 <sup>6</sup>	3 2.74 6.10 <sup>-1</sup> 3.35 = 10 <sup>-1</sup> 2.36 6.10 <sup>6</sup>	4 9.79 6.16** 1.28 6.10** 2.16 6.18*	3 1.23 n 10° 1 2.73 n 10° 1 1.1- n 10°	0 2.74 4.10-1 3.53 1.10-1 1.10 1.0*	7 2.74 m 10 <sup>-1</sup> 3.53 m 10 <sup>-6</sup> 1.16 m 10 <sup>0</sup>	8 8.7¢ n 10-6 8.61 n 10-6 3.83 n 186	3 1.33 6.36 <sup>-1</sup> 1.73 m 10 <sup>-1</sup> -,73 m 10 <sup>-6</sup>	10 2,76 6,10°1 3,30 6,10°1 3,99 8,10°3	ti 2,76 m 10° b 1.06 - 10° 5,99 m 10°	12 8,78 8,10° 6 8,61 8,10° 6 2,63 6,10°	13 3,70 m 10 <sup>-1</sup> 3,76 m 10 <sup>-1</sup> 3,06 m 10 <sup>-6</sup>	16 2.76 0.10-1 1.06 0.10 <sup>0</sup> 3.04 0.12 <sup>0</sup>	15 3.53 x 10 <sup>-2</sup> 6.59 x 10 <sup>-6</sup> 1.39 x 10 <sup>-8</sup>	16 3, 53 110° 2 3, 42 110° 2 1, 51 110°
(÷1)	y*= 2010 mm*1	e1+3856 mm <sup>-1</sup>	y'=28% m*1	v*= 2878	V*= 2000 mg* t	A.= 5608	v*= 2000 mm*	40 5000	V*x 37000 CH**	y*= 2800	>1= 2000 MI <sup>11</sup>	v*= 2790 mp*1	v1= 2700	· *= 2700	y*= 2700 mm**	1°=3700 cm° 1
2856.00 2845.00 2840.60 2815.80 2836.83	0. F. 0.	0. 0. 0. 0.	0. 0. 0.	0. 0. 0.												
2623.00 2629.00 2715.90 2619.05 2605.00	0. 6.001 0.034 0.034	0. 0.004 0.014 8.039	0 0. 0. 0.	0. 0. 0. 0.												
2000.00 2795.00 2790.00 2745.00 2780.00	0.299 0.451 0.7-2 0.945 1.313	C.125 C.221 O.315 O.444 O.454	0.000 115 0.102 0.179	0.000 8.022 C.064 C.107 3.176	0. 0. 0. 0.	0. 0. 0.00. 0.01:	0. 0. 0.000	0. 0. 0. 0. 785	0. 0. 0. 0.	0. 0. 0. 0. 0.001	0. 0. 0.					
2715.00 2110.00 2745.00 2745.00 2745.00	1.952 2.697 3.311 4.165 4.662	0.954 1.746 1.786 2.149 2.313	0.3G1 0.474 0.473 0.662 0.432	0.498 0.488 0.670 0.628 077	0.015	0.040 0.124 0.211 0.280 0.294	0.044 0.114 0.204 0.205	0.028 0.04 0.104 0.147	0.02- 0 031 0.032 0.122	0.025 0.042 8.104 0.132	0.045 8.090 0.137 0.172 0.161					
2750.00 2745.00 2740.09 2735.00 2 110.00	9,821 9,999 6,104 6,890 7,179	2.119 2.927 3.:11 1.640 3.622	5.91a 1.000 1.211 1.382	0.961 1.110 1.111 1.449 1.11*	8.237 0.319 0.408 0.419 0.914	0.323 0.166 0.665 0.523	0.713 0.404 0.488 0.554 0.595	0.167 9.225 0.274 0.325 0.349	0.140 0.178 0.219 0.240	0.144 0.204 0.233 0.289 0.301	0.204 0.244 0.204 0.148 0.571					
2125.00 2720.00 2115.00 2110.00	7.34 1 7.563 7.461 7.843 2.144	0.025 0.010 0.112 6.210 0.170	1.429 1.429 1.429 1.429	1.536 1.567 1.672 1.673 1.741	0.321 0.321 0.321 0.321	0.334 0.934 0.334 0.334	0.606	0.352	0.287 0.287 0.267 0.267	0.101 0.301 0.101 0.101	0.373 0.373 0.873 0.373 0.374					
2700.80 2995.00 2490.00 2485.00 2440.00	6.46; 6.490 9.191 9.900 10.763	4.372 4.764 5.024 3.126 3.426	1.424 1.424 1.424 1.424	1.622 1.125 2.256 2.216 2. 96	0.521	0.358 0.358 0.369 0.362 0.463	0.604 0.604 0.622 0.637	0.352 8.352 0.352 0.352	0.28* 0.287 0.267 0.287	0.101 0.101 0.301 0.308	0.313 0.373 0.361 0.404 0.426	o. o. o.	o. o.	0. 0. 0.	#. 0. 0.	0. 0. 0.
2073.00 _070.00 2003.00 2340.00 2055.00	11.514 12.493 13.464 14.141 15.457	A.111 9.700 7.039 7.607 F.149	1.463 3.500 1.622 1.734 1.496	2.617 2.663 1.119 3.176 1.661	0.549	0.429 0.444 0.703 0.734 0.478	0.473 0.701 0.743 0.749 0.446	0.312 0.732 0.160 0.168	0.281 0.292 0.320 0.357	0.134 0.134 0.181 0.411	0.463 0.504 0.546 0.547	8. 0. 0.	0. 0. 0. 0.001	0. 0.002 0.004	0. 0. 0.	c. o. o.
2050.00 2050.00 2000.00 2010.00	16.737 18.376 20.636 23.676 27.167	0.6.4 7.619 11.162 11.141 15.531	2.144 2.110 1.141 4.194	4.016 4.016 154 445 7.947	0.610 0.958 1.216 1.630 2.171	0.919 1.149 1.5G6 2.042 2.747	0.977 1.157 1.492 2.044 2.474	0.492 0.577 0.713 0.928	0.447 0.550 0.491 0.930 1.251	0.401 0.792 1.084	0.742 0.665 1.078 1.425 1.425	0.006 0.016 0.100 0.217 0.110	0.041 0.084 0.144 0.304	1 .054 0 114 0. 17 2-1 2	0.004 0.012 0.081 0.151	0. 0.004 0.025 0.047 0.140
2625.90 2620.30 2615.00 2615.00 2605.00	10.709 31.666 15.490 16.412 41.614	17.664 14.866 21.026 22.659 24.917	6.4 *2 f. 13 6 f. 7; 7 3.44;	4.45a 10.4.2 11.22a 12.144 13.449	2.713 3.216 3.461 3.719	1.444 4.02. 4.266 4.611	3.750 4.418 4.818 5.240 5.343	1.493 1.751 1.861 1.020 2.219	1.010 1.945 2.140 2.268 2.547	1.697 2.212 2.137 4.540 2.865	2.753 2.727 2.915 3.149	0.535 0.671 0.736 0.791	0.131 0.928 1.016 1.049	0.411 1.072 1.157 1.283	0.214 0.272 0.294 0.313 0.333	0.237 0.306 0.332 0.336 0.416
2600.00 25.5.03 2590.00 2563.00 2563.00	43.140 48.194 51.211 32.616 55.672	27.271 29.344 31.67m 12.444 33.605	10.51* 11.469 17.211 12.745 1120	1 120 10.183 17-161 17.349 28.546	4.7/2 1.167 5.541 3.775 5.924	3.934 6.302 6.934 7.222 7.914	6.670 7.116 8.136 8.595 =.721	2.490 2.724 2.915 3.016 3.117	2.653 6.162 3.408 3.576	3.235 2.335 3.796 3.947	1.990 4.466 4.710 4.956 5.131	1.049 1.172 1.271 1.334	1.443 1.672 1.616 1.915	1.722 1.912 2.001 2.219 2.7 8	0.403 0.433 0.500 0.319 0.519	0.494 0.333 0.339 0.620 0.620
1575.07 170.00 165.00 2560.03 2155J	57.564 ** 416 5368 61.472 65.840	14.656 13.745 34.763 37.471 37.345	: .146 33 533 13-5 7 11.964 14.224	14-127 14-465 20,141 20,904 21-441	0.019 0.006 0.132 0.177 0.229	7,343 7,456 7,752 7,650 7,969	4.194 9.455 4.720 9.467 11.322	3.176 9.202 9.215 1.265 9.270	6, 746 3, 768 3, 023 3, 838 3, 891	4.116 4.194 4.209 4.241 4.244	5.271 5.414 5.367 4.135 4.915	1.402 1.910 1.422 1.437	2.031 2.041 2.048 2.101 2.142	2.* 2. 2.574 2.672 2.769	0.519 0.519 0.519 0.519 0.513	0.632 7.632 0.632 0.632 0.632
2550.00 2545.00 2540.00 2515.00 2510.00	68.437 F1.119 F4.469 FF.944 61.557	40.869 42.601 44.535 48.710 49.217	14.514 14.662 15.275 14.754 14.174	22.50e 23.686 24.5e7 25.656 27.161	6.101 6.105 6.106 6.651 6.640	6.*10 679 6.662 6.79( 9.061	10.700 11.121 1. 445 12.155 12.127	1.361 1.140 3.549 1.444 3.512	1.927 3.566 4.02 4.011	4,364 4,444 4,540 656 4, 75	4.124 6.163 6.654 6.679 7.360	1.477 1.500 1.510 1.571 3.621	2.196 2.260 2.394 2.424 2.31	2.426 3.077 3.233 3.425 3.664	0.514 0.515 0.517 0.721 0.731	0.612 0.612 0.612 0.642 0.664
2525.00 2525.70 2515.00 2110.30 2105.00	85.444 69.959 94.574 99.236 104.033	92.092 99.310 94.924 62.744 4-234	10.1ml 10.1ml 10.415 20.647 22.323	29.694 11-101 11-12 10-694 39-692	7 047 7.446 7.962 6.432 8.477	10 7 10.767 11.601 12.412	1 .052 10.695 .5.638 17.235 30.617	3.650 3.665 4.064 6.248 4.884	4.327 4.513 4.764 3.097 5.301	5.007 5.201 5.630 4.064 6.324	7,*70 6.367 9.644 4.649 10.996	1.690 1.790 1.945 2.076 2.227	2.662 2.611 1.036 1.306 1.364	3.914 4.235 4.620 3.056 3.314	0.344 0.467 0.657 0.714 0.744	5.696 5.766 0.615 0.666 0.978
2500.00 2493.00 2490.00 2444.00 2460.00	104.6+7 111.754 116.746 123.486 126.6+1	70.714 75.257 74.865 84.613 89.444	23.475 27.194 27.124 23.376 11.482	42.049 41.274 49.890 92.994 97.207	9.417 2.646 10.511 11.364 12.181	19.166 14.926 13.144 14.334 16.741	20.204 21.814 23.613 26.294 28.627	4.641 4.636 3.034 3.183 3.701	3.633 3.674 6.203 4.614 7.073	4.934 T.334 F.906 4.511 9.410	11.332 12.460 17.399 14.474 16.434	2.342 2.431 2.363 2.736 2.930	1.741 4.024 4.146 4.717 3.171	5.763 4.494 736 7.891 8.743	0.797 0.800 0.612 0.617 0.127	1.044 1.044 1.144 1.274 1.392
2473.00 2410.00 2443.00 2443.00 2433.00	71.6+1   110.661   41   7   146.681   51.4#1	44.351 44.105 164.103 144.103	79.619 37.397 60.845 44.537 46.663	41.464 46.741 7' 14. 64.3 44.7	13.132 14.160 13.313 16.646 14.332	19.702 21.964 23.664 24.196 24.210	11.395 34.439 17.749 41.460 43.333	0.033 0.373 0.727 7.140 7.079	7.949 8.941 8.969 9.401 13.310	10.297 11.262 12.425 13.192 13.512	18.115 19.966 22.110 26.518 27.361	1.044 3.746 41 .666 3.967	5.661 6.828 4.589	9,703 10,797 2,031 13,336 13,756	0.973 1.006 1.034 1.033 1.311	1.515 1.644 1.791 1.975 2.226
2440.00 2441.00 2440.00 2411.00 2410.00	150.023 163.467 166.611 173.647 176.641	174.371 124.371 124.101 154.301 153.503	57.161 57.904 62.004 67.769 77.762	85.970 90.946 90.446 100.476	20.574 21.291 21.407 73.145 34.231	\$7.770 34.779 41.163 45.44 30.704	30.003 34.713 34.574 44.510 65.478	0.447 9.470 10.*95 12.314 14.119	11.004 13.2.7 15.328 17.765 20.666	17.741 20.461 23.731 27.457 31.371	10.464 34.447 18.145 43.163 47.689	4.422 3.044 3.821 4.744 7.423	11.444 13.128 14.400 14.340	17.609 20.129 23.499 27.088 31.007	1.241 1.42 1.724 2.043 2.393	7.624 3.192 3.919 4.764 5.696
2475 ; 2420.60 2415.65 2410.60 2405.00	163.627 166.681 193.663 146.621 263.641	1.4.101 402.103 154.101 144.101	77.211 82.712 11.107 52.762 17.782	130.944 117.335 125.944 125.944	18.616 43.417 46.111 11.276 21.277	55 669 60.623 65.622 70.622 73.622	74.4°7 79.472 64.672 69.6°2 69.672	10.100 19.012 22.119 26.139 30.412	21.956 27.813 32.185 36.811 41.720	16.0-8 40.822 45,719 50.711 55.703	92.74. 97.754 62.444 47.179 72.679	9.129 10.876 .1.CAO 15.750 19.004	22.456 26.678 11.365 36.171 41.005	38.372 40.727 44.461 49.621 34.795	2 806 * 431 4 161 5 161 6 470	n.760 6-211 10-078 12-696 15-227
2400.00 2195.00 2196.03 2175.00 2363.00	205,441 213,453 218,453 221,413 221,413	144.101 174.101 179.101 164.101	107.7n2 107.7n2 112.7n2 117.7n2 117.7n2	195.00A 105.336 105.00A 190.00A 155.30A	61.273 66.273 72.277 76.2°7 61.273	39.422 85.427 90.422 45.622 100.072	99.472 194.377 199.472 114.472 119.472	15.047 14.931 44.846 49.672 54.871	46.473 51.4.6 56.838 61.664 60.668	60.753 65 03 70.703 73.703 60.703	77.474 92.479 67.679 92.119	22.042 27.228 32.023 36.931 41.960	43.471 50.457 13.452 40.457 63.152	\$9.747 64.743 69.743 74.793 19.193	0.166 10.416 11.461 10.346 26.329	10.707 22.034 27.391 12.570 17.366
2375, 30 2170, nr 2+03+30 2+03+00 2315+00	211.009 210.001 241.502 246.008 253.003	196,308 356,171 226,1 1 203,17 216,203	127 742 112.762 117.742 14 .762 .7.76.	185.34A 165.328 177.346 179.368 186.944	46.211 91.213 46.171 161.211 10277	.05.627 110.622 119.622 120.622 125.672	124.472 129.47* 114.4 2 119.472	59.671 64.67* 69.371 ) 471 F3.671	71.005 70.000 81.008 80.000 91.005	95.763 90.703 93.703 100.70* 105.703	152.6 . 172.676 112.679 1.7.679	\$6.960 \$1.960 \$6.960 \$1.960	70.452 13.952 60.952 83.952 90.952	64.763 69.791 90.793 90.793 104.793	10.124 11.121 18.121 41.121 16.121	42.364 47.369 32.346 57.564 62.346
2395.05 2395.05 2800.00 2319.70 2310.02	258.4+5 268.4+5 744.45 - 278.6+1 278.6+1	214.1 1 227.1 1 227.4 1 227.61 21.131 17.301	152.762 157.752 167.762 167.752 172.762	18".47 h 197.4 ; 141.° ; 200.' Fa 205. 446	117.273 1.m.2 1 123.273 124.273 161.271	130.622 139.472 140.622 145.622 130.422	.49.672 134.472 139.472 144.47* 169.41.	84.673 89.675 94.871 99.871 394.871	97.440 101.000 100.440 111.440	117.701 115.703 127.703 125.703 130.703	127.474 132.474 137.41+ 142.474 147,474	009.41 009.46 009.60 007.00	63.452 160.452 163.452 710.452 115.44	104.743 114.743 114.743 124.741 127.743	51.121 58.141 61.723 66.121 73.123	67.566 77.566 77.566 62.566 67.564

				Ta	bre	4-1		· (ν)	${\sf d}  u$	(c	ont'c	1)				
Sam, Mo, p(ntm) P <sub>e</sub> (ntm) u(ste cm) STP	3 5.76 m lan <sup>1</sup> 2.75 m lon <sup>1</sup> 8.44 m lof	2 5 76 = 10 <sup>-5</sup> 2.75 = 10 <sup>-2</sup> 4.24 = 10 <sup>6</sup>	3 2, F - 4, 12 - 1 2, 36 8, 10 - 1 2, 36 8, 10 - 1	9,79 h 10 <sup>1</sup> l 1,28 m 10 <sup>6</sup> 2,14 m 10 <sup>4</sup>	5 1,37 ± 10** 1.73 ± 10** 1.14 ± 10**	9 2.74 ± 10° 1 3.56 ± 10° 5 1.19 ± 10°	7 2.74 n 10°1 5.97 n 10° P 1.15 n 10°	6,74 10 <sup>-3</sup> 6,63 10 <sup>-2</sup> 5,63 10 <sup>6</sup>	5 1.73 4.10°5 1.73 a.10°7 5.75 a.108	10 2, 7a 4 10" f 3, 36 x 10" f 3, 99 4 10 <sup>6</sup>	11 2:74 10 <sup>-1</sup> .:06 n 10 <sup>0</sup> 5:99 n 10 <sup>9</sup>	12 6, 78 8 10 <sup>-8</sup> 6, 61 8 10 <sup>-8</sup> 2, 97 10 <sup>6</sup>	17 2.74 a 10 <sup>-1</sup> 3.36 a 10 <sup>-1</sup> 3.06 n 10 <sup>8</sup>	1.74 n 10°1 1.06 n 10°2 3.06 n 10°2	57 3, 53 n 10°8 4, 50 n 10°8 7, 33 n 10°	16 3,33 a 19-8 3,62 a 10 <sup>-1</sup> 1,33 a 10 <sup>8</sup>
(cm' <sup>(</sup> )	1 = 2450 √m² <sup>1</sup>	_1+ 2850 cm*1	* 2850 cm 1	214 78 CBT	2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	v1+ 2 #00 cm <sup>-1</sup>	v*+ 2800 ma* 1	1 <sup>3</sup> • 2800 cm <sup>-1</sup>	."+ 2 <b>800</b> cm"."	v'=7800 cm <sup>-1</sup>	չ'⊾ 2800 eni⁻²	2700	,1-2100 cm <sup>-1</sup>	y) a 2700	ν•ω , ~,0 cm <sup>- l</sup>	1° = 2700 40° 1
2324.00 2324.00 2335.00 2310.00 2305.00	201-093 201-083 291-083 298.083 501.081	244.301 244.301 254.303 214.303 764.301	1 ff. 7a2 1 a2. 7a2 1 a7. 7a2 1 a2. 742 1 97. 762	215.966 215.966 220.966 2/5.986 2/6.968	138.273 141.273 146.273 157.273 156.277	155.622 163.622 165.622 170.622 175.622	174.472 179.472 189.472 189.472	140.071 119.071 124.671 124.671	121.500 129.800 131.668 130.668 141.648	135.707 145.701 150.701 150.703	157.679 157.679 167.679 167.979 172.976	96.980 101.980 108.940 111.980 116.980	120.952 125.962 130.952 135.952 140.952	134.793 139.793 196.793 149.793 154.791	78.123 61.123 68.325 93.321 96.121	92.596 97.586 102.566 107.568 112.566
23+0.00 224-0.00 2245-00 2245-00	108.623 113.493 716.623 163.881 326.663	284.303 274.303 274.303 284.303 203.4-1	207.762 207.762 212.762 217.762 222.762	235.900 240.900 245.900 250.960 254.000	141.271 140.273 177.273 174.273 163.277	180.622 185. 32 190.622 195.622	199.472 204.472 209.412 234.472 (19,472	134.871 137.871 144.871 144.671 144.873	144.468 151.468 150.448 181.468	180.703 165.703 170.703 175.703	177.679 162.479 167.479 1*2.679 197 19	121.980 174.980 131.980 138.980 141.940	145.952 150.992 135.452 180.752 38652	159.793 144.793 169.793 174.795 379,793	101.323 107.321 115.323 116.323 124.323	117.566 122.566 122.588 172.568 172.568
2275.00 2270.00 2265.00 2266.00 2255.00	113.663 138.661 143.661 146.663 513.463	294. i 299. tu3 3u4. i03 107. i03 314. iu1	227.142 232.162 231.142 242.752 247.162	780.966 265.966 777.966 275.968	186.273 145.273 198.273 269.279 268.277	209.622 215.622 215.622 229.622	224.472 224.482 214.472 219.482 244.472	159.8*1 164.671 169.871 174.671 178.671	171.008 170.008 181.008 180.008	185.703 190.763 195.701 200.703 205.103	202 9 26 - 79 212-679 217-679 222-879	148.980 151.980 156.980 163.980 167.980	170.452 175.452 180.454 185.452 140.452	184.793 189.793 184.793 194.793 204.793	128.323 176.323 136.323 141.321 148.720	142.548 147.548 152.548 157.588 142.584
2250.00 2245.00 2244.60 2215.00 2210.60	171,661 161-263 166-823 173-663	317.301 324.303 379.30. 334.3u3	252.1e2 257.7e2 262.162 261.1e2 272.7e2	281.416 295.518 295.525 300.966 305.966	713.275 210.277 227.271 224.273 213.273	230.622 235.627 240.622 245.627 250.422	244.472 254.472 259.472 264.472 299.472	134.871 1/3.871 1/4.871 1/4.848 204.776	196.663 201.668 206.668 217.668 210.865	210.701 215.703 220.703 225.701 230.701	227.678 252.679 217.679 242.629 247.619	173.740 178.973 181.996 188.776 191.380	195.932 200.992 205.952 210.943 215.023	209.16° 214 ~43 9.793 224.793 229.792	153.302 :38.151 !82.818 386.953 170.490	167.588 172.588 177.566 182.517 187.375
2225,00 2226,66 2715,00 2216,00 2205,00	783.647 384.643 398.683 403.683	36+.3-3 345,1,1 164.351 159.303 364.323	277.762 292.762 261.760 292.661 297.272	110.466 135.468 120.966 125.183 310.914	214.250 241.181 247.861 251.270 255.359	255.622 266.612 265.653 269.859 273.432	274.472 274.412 254.472 265.440 274.094	209.493 213.701 237.242 219.699 221.650	221.50 224.285 230.520 233.941 236.533	249.481 249.401 245.162 249.084 252.208	252.476 257.855 242.551 267.060 270.894	195.550 196.785 201.703 201.619 204.958	220.787 225.313 229.169 232.205 239.360	234.782 239.888 244.157 24f.846 250.608	173.576 175.649 177.570 176.749 179.541	181.840 195.559 198.574 208.325 201.590
.200.00 7195.00 2196.00 2185.00 2186.00	408.669 413.669 414.879 423.643 428.5/9	369.29G 374.221 379.029 383.573 367.714	101.1.2 304.479 307.671 109.877 311.634	115.037 740.47d 344.569 348.20 351.151	254.0[8 260.063 241.646 252.748 263.586	270.675 278.065 260.851 482.396 253.215	297.061 50.412 504.034 505.363 507.501	23.250 224.244 224.842 225.453 225.778	239.941 240.951 241.668 242.181	2*4.57f 258.234 257.457 258.318 756.350	279.684 278.168 277.840 278.648 279.953	205 254 206.479 206.894 203.192 207.391	235.858 236.871 237.597 238.046 238.387	252,536 253,862 254,746 255,788 255,956	180.053 160.6G0 180.847 160.777 180.661	202,388 402,905 203,234 203,433 203,549
2175.00 2176.00 2165.00 2161.00 2355.00	437.250 417.761 442.219 448.803 491.246	391.478 394.814 396.316 461.846 465.177	315.0CB 314.125 315.488 *17.445 318.745	354.007 356.798 356.796 461.317 363.569	264.230 244.725 225.823 266.492 267.226	281, 931 284-657 265-915 266-612 287-762	209,732 309,743 311,791 312,438 313,438	228.029 228.233 226.751 223.118 227.462	242.982 247.891 241.458 244.011 244.495	258.420 259.793 260.473 261.083 263.608	280.994 289.228 292.121 282.864 289.675	207.539 207.649 207.611 208.194 208.392	230.638 230.637 238.265 239.606 241.902	256.224 256.519 257.071 257.493 257.462	190.919 190.955 161.093 161.200 161.274	203.615 203.654 203.787 203.614 203.862
2100.00 -145.41 -2141.10 -2141.10 -110.10	455.634 458.884 484.235 488.795 471.544	462.445 411.422 415.053 412.745 422.469	120,404 121,253 327,565 325,563 326,129	345.770 030133 170.00 170.00 170.00 170.00 110.00	267,950 268,715 269,615 270,857 272,364	282,694 288,642 290,917 292, 15 294,656	3:5:206 3:6:498 3:7:966 3:7:743 4:7:527	227.808 278.165 278.626 279.150 210.161	244.959 245.454 246.0"1 246.7 246.8	202.540 202.684 203.702 284.109 205.505	284.337 285.123 285.923 286.977 2-0.358	208.599 204.827 204.169 209.453 210.410	2+0-202 2+0-516 2+0-905 2+1-392 2+2-618	258.231 258.023 259.076 258.633 267.023	161.355 161.449 161.512 161.731 162.230	203.907 263.949 264.302 204.128 204.915
21/h.cu 21/1.co 21/1.co 21/h.t6 21/h.co	410.496 483.473 486.472 493.477 436.477	432 502 437.452 437.452 442.451	111.000 134.71, 138.510 143.132	700-346 3 +01 5 - 91 374-74, 345-656	274,430 210,492 218,854 261,602 284,848	296.975 299.485 702.621 308.409 309.954	3/5.835 3/9.482 131.038 138.114 342.477	2:1:972 2:32:647 2:31:979 2:35:7:8 2:37:564	288.699 251.019 252.763 254.012 257.324	267.503 269.352 271.645 274.412 277.220	291.247 293.820 296.990 500.70 304.299	211.399 212.175 213.132 2:4.386 21~ .94	243.99* 245.28: 246.945 248.989 291.052	282,878 264,645 2 6,974 26, 749 272,518	182,978 183,766 183,865 184,566 184,587	265.944 206.747 267.809 209.390 210.949
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7 Filts In Thint Thatin Thatin Institut	5.6.472 533.472 536.472 583.472 546.472	677,657 52,551 667,651 69,451	474.72m 174.72m 144.72m 18.72m 18.72m	428.90; 413.50; 432.52; 443.50; 443.50;	301.524 309.81 314.215 314.645 123.079	334.250 318.067 343.909 347.777 753.647	970,264 37°,286 38.,266 385,264 390,264	251.6%1 254.812 257.325 266.5 0 263.5 %	274.653 278.044 .81.775 285.530 238.276	297.540 709.884 309.118 310.791 315.288	328,541 337,473 334,410 343,37; 348,135	227.134 229.238 291.546 233.659 298.130	287.381 27'.970 27 .722 278.489 282.248	292.130 285.851 301.479 304.201 110.901	192.663 171.763 190.095 196.411 197.666	222.499 225.100 477.968 250.768 231.452
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#### SECTION 5

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Absorption by CO <sub>2</sub> Between 1800 and	2850 cm <sup>-1</sup> (3,5-	-5.6 Mic	erons)
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#### 3 ABSTRACT

Transmission spectra in the 1800-2850 cm<sup>-1</sup> region have been obtained for more than 100 samples of CO<sub>2</sub> and CO<sub>2</sub> mixed with N<sub>2</sub> and A. The spectral resolution was 2.5 cm<sup>-1</sup>. Sample pressures varied from 0.0055 to 742 torr with absorber thicknesses covering the range from 0.081 to 84,400 atm cm<sub>2</sub>. Spectra of several samples at the lower pressures show the effect of Doppler broadening. Measurements in the 2400-2560 cm<sup>-1</sup> region provide information about the absorption by the extreme wings of collision-broadened lines. Replotted transmission spectra and extensive tables of integrated absorptance for 116 samples are included.

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Security Classification ta. LINK A LINK B LINK C KEY WORDS ROLE WT ROLE WT ROLE WT CO2 Infrared Absorption Doppler Broadening Collision Broadening

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